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Cosmic inflation with inhomogeneous initial conditions

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We investigate the robustness of inflation models to inhomogeneous initial conditions. In the simplest case inflation is described by a single scalar field ϕ minimally coupled to gravity with homogeneous initial value ϕ_0 and canonical momentum $\Pi_0 = 0$. However, there is no fundamental reason to reckon that inflation started with such homogeneous decomposition to study fast-roll initial conditions i.e., nontrivial ϕ_0 and Π_0 for attractor (small-field model) and Starobinsky (large-field). We find that for small field inflation, with initial inhomogeneous scalar field profile inflating bubbles while for non-homogeneous evolution depends on the relative phase between initial scalar field profile and ϕ_0 confirming previous results that large field model is more robust than small field one.

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