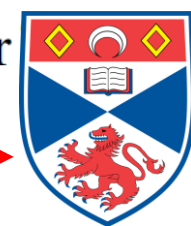


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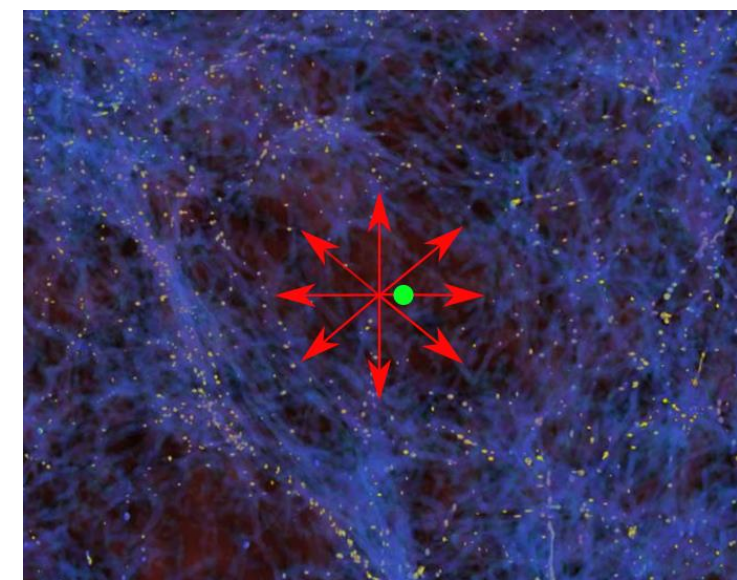
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MNRAS, 527, 4388 – 4396 (January 2024)

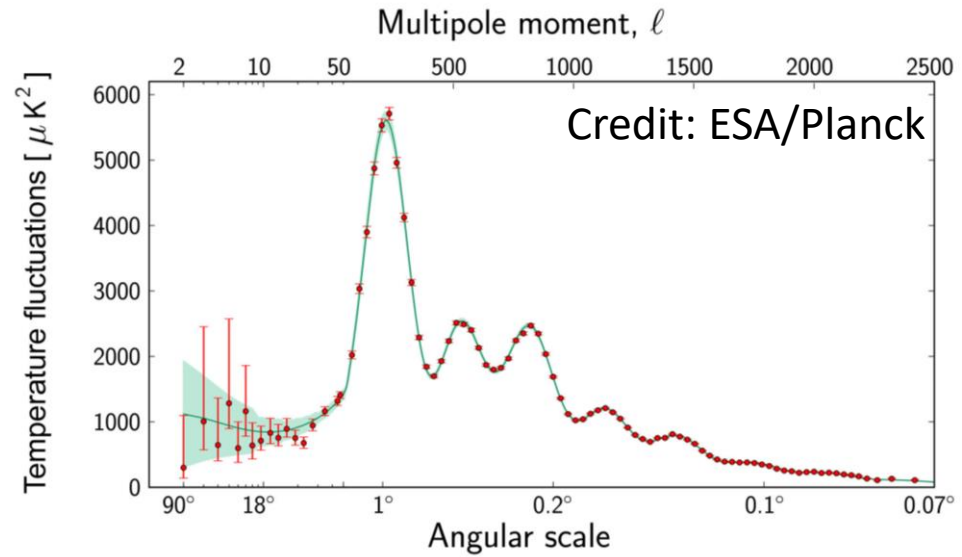
Joint solution to the Hubble and bulk flow tensions using the observed local supervoid and faster structure growth



Two routes to the present expansion rate

- The currently popular Λ CDM model has achieved many successes (e.g. primordial D and He, galaxy cluster mass function at low z , cosmic shear)
- Its parameters are usually calibrated against CMB anisotropies ($z = 1100$)
- Λ CDM then predicts the present expansion rate

$$H_0 \equiv \dot{a}/a$$



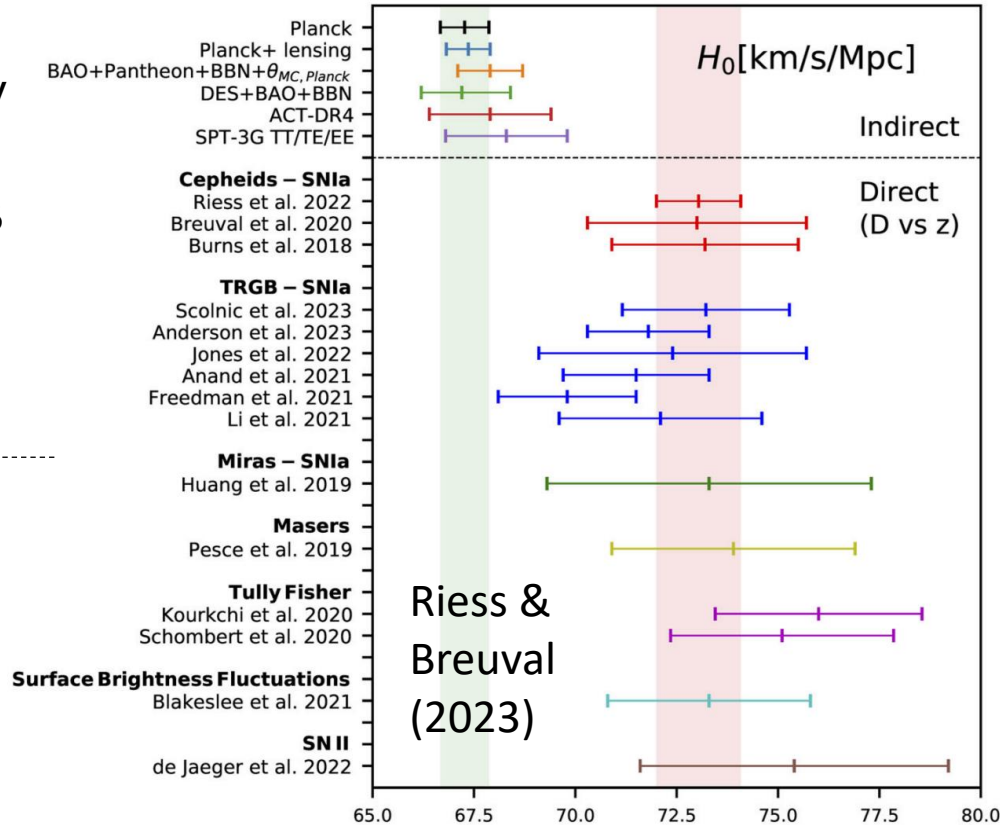
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- Local observations show a nearly linear relation between distance and spectroscopic redshift z
- If this is assigned entirely to the difference in a at the time of emission and reception, we can infer

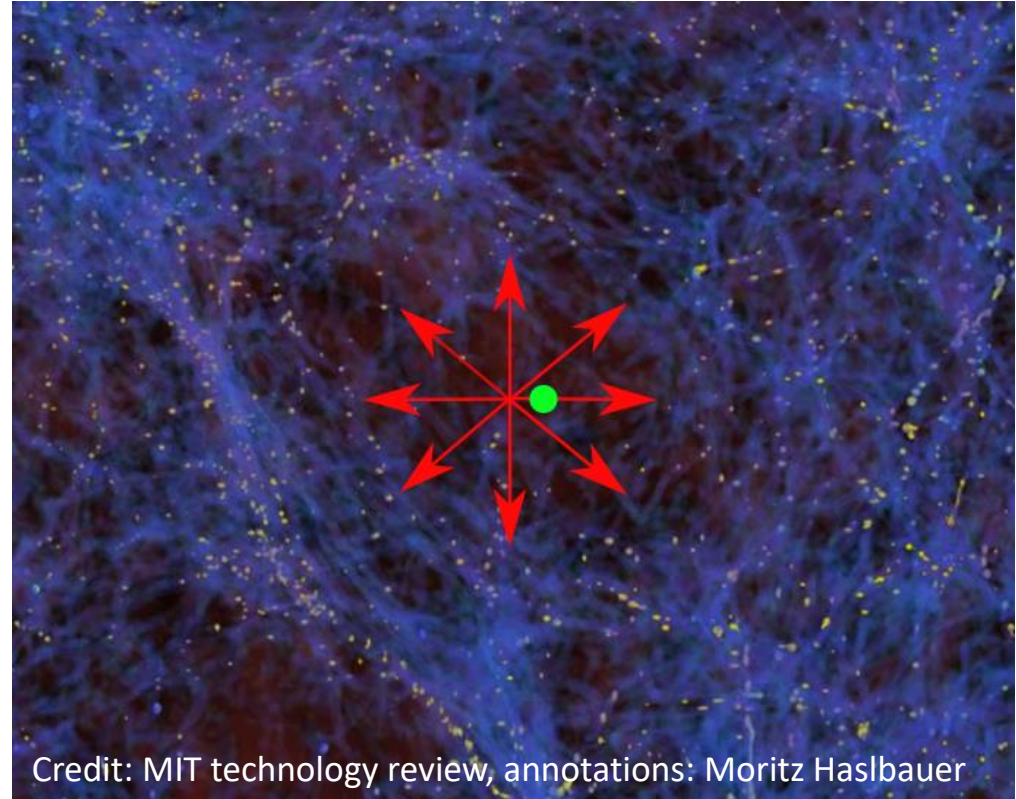
$$H_0 = \lim_{z \rightarrow 0} c \frac{dz}{dr}$$



The impact of local structure

- The local measurement is affected by peculiar velocities, which are velocities in the CMB rest frame (e.g. M31 is approaching us)
 - As more distant galaxies are considered, the impact of peculiar velocities should decrease
 - But if our location (green dot) is near the centre of a large void, they could rise at first and only start decreasing quite far out
- Peculiar velocities might skew local H_0 .

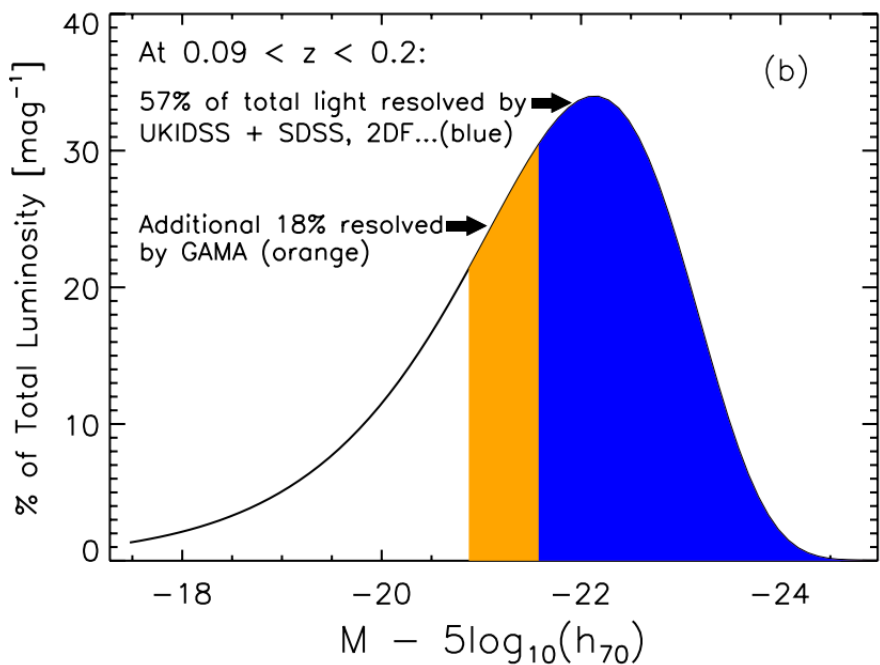
$$H_0 = \lim_{z \rightarrow 0 \text{ (but not too low)}} c \frac{dz}{dr}$$



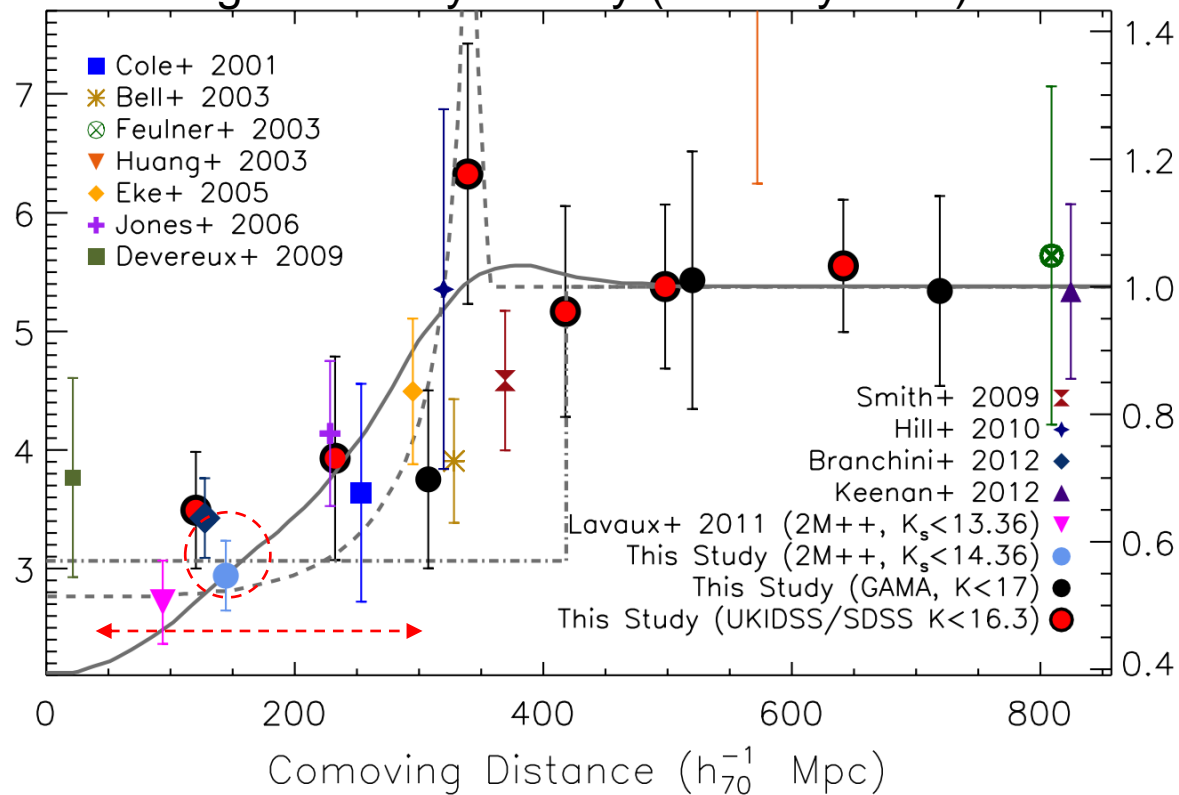
Credit: MIT technology review, annotations: Moritz Haslbauer

The KBC void (Keenan+ 2013)

- Galaxy number counts over 90% of sky and most of the luminosity function show an underdensity at 40 – 300 Mpc.

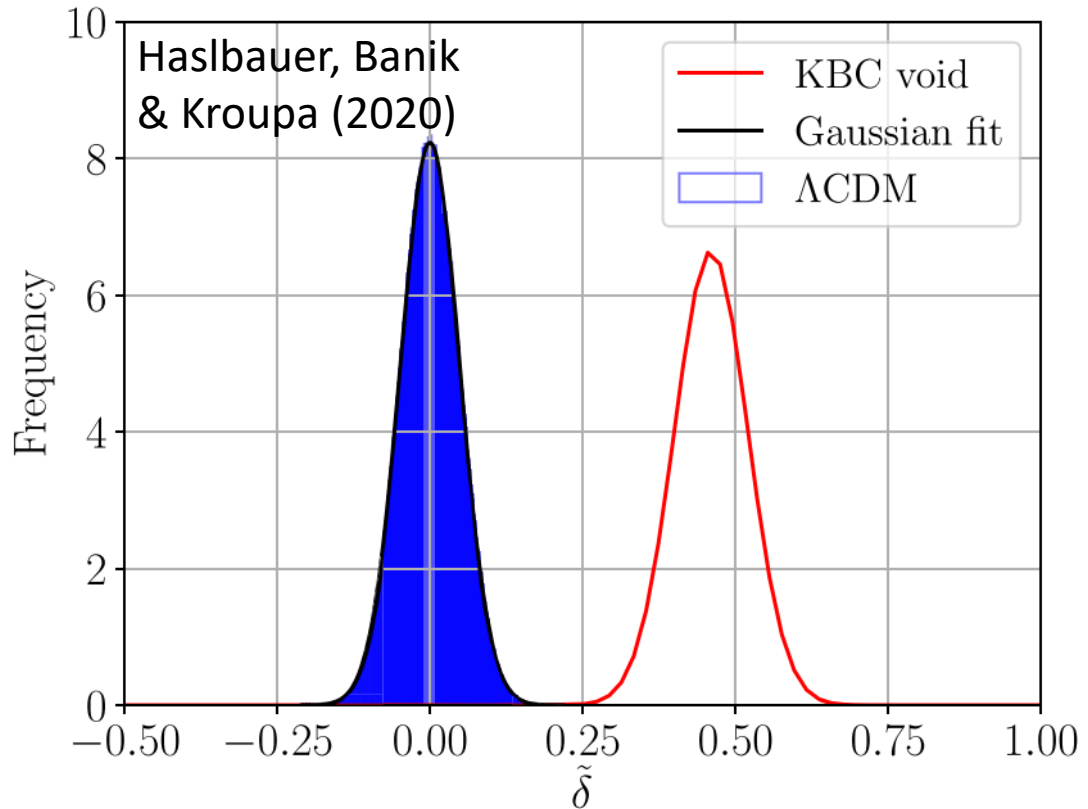


Comoving luminosity density (arbitrary units)



Analogues in Λ CDM

- Uniform grid of 10^6 vantage points in Millennium XXL (Angulo+ 2012), a Λ CDM N -body simulation with 4.1 cGpc box size.
 - Get total mass in halos with semi-analytic $M_* > 10^{10} M_\odot/h$ within 40 – 300 Mpc of each vantage point
 - Compare density to cosmic mean value
 - Enhance the density contrast 1.5x to allow for redshift space distortions: observers think they are seeing out to some distance d based on the redshift z , but outflows from local void mean the actual distance $< d$, reducing the galaxy number count
- Tension with Λ CDM is 6.04σ .



Increasing the cosmic variance

- Cosmic variance in local measurements of H_0 should be quite small in Λ CDM

➤ Cannot solve the Hubble tension

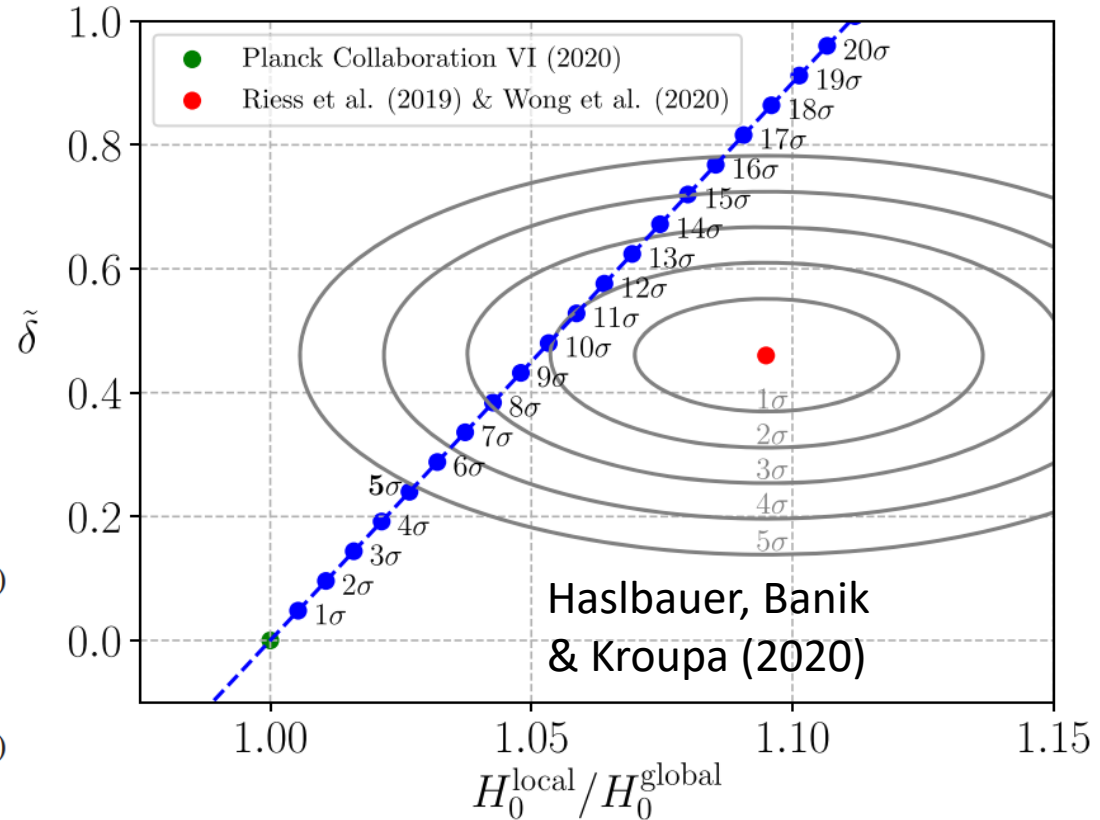
- But structure formation must be enhanced to explain the KBC void (local void also found by Wong+ 2022)

- Can we relate the observed local void to the Hubble tension?

$$\frac{\Delta H}{H} \equiv f\delta, \quad (18)$$

where e.g. Marra et al. (2013) showed that for $\delta \ll 1$ in Λ CDM,

$$f = \frac{\Omega_m^{0.6}}{3b}, \quad (19)$$



MOND simulation of outflow from void

- Start with small initial underdensity at $a = 0.1$, when the expansion rate = H_i

$$\ddot{\mathbf{r}} = \mathbf{g}_{\text{void}} + \overbrace{\frac{\ddot{a}}{a} \mathbf{r}}^{g_{\text{Hubble}}},$$

$$\dot{\mathbf{r}}_i = H_i \mathbf{r}_i,$$

- Acceleration beyond the cosmic expansion term g_{Hubble} is due to mass deficit interior to r
- Need present cosmic mean matter density ρ_0

$$g_{\text{N}} \equiv \frac{G \Delta M}{r^2}, \quad \text{with}$$

$$\Delta M \equiv \frac{4\pi}{3} \rho_0 \left(\frac{r}{a}\right)^3 - M_{\text{enc}},$$

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- To get enhanced structure formation, the gravity is enhanced as in MOND (as proxy for unknown model)
- External field weakens this enhancement and causes void to move as a whole
- Important for velocities in CMB frame.

$$g = g_{\text{N}} \left(\frac{1}{2} + \sqrt{\frac{1}{4} + a_0 \left(g_{\text{N}}^2 + g_{\text{N,ext}}^2 \right)^{-\frac{1}{2}}} \right)$$



Enhanced structure growth

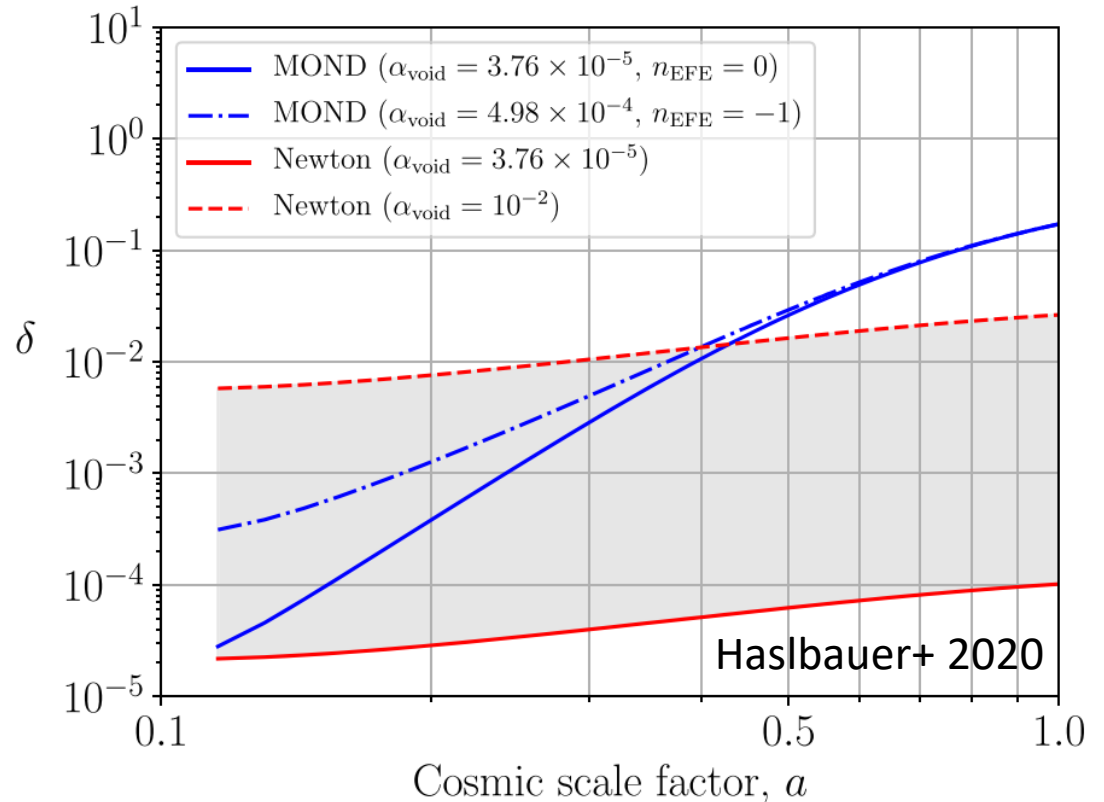
- Figure: evolution of density contrast on 300 cMpc scale

- Normally, matter-dominated era has

$$\delta \propto a$$

- Structure formation about 4x faster in MOND

- Same δ today associated with larger peculiar velocity and thus larger enhancement to H_0 .



Redshifts are not purely cosmological

- Model includes three sources of redshift:

- Cosmic expansion
- Peculiar velocity away from void
- Gravitational redshift

$$\frac{\lambda_{\text{obs}}}{\lambda_{\text{emit}}} = \frac{1}{a(t)} \overbrace{\sqrt{\frac{c + v_{\text{int}}}{c - v_{\text{int}}}}}^{\text{Doppler}} \overbrace{\exp\left(\frac{1}{c^2} \int g_{\text{void}} dr\right)}^{\text{GR}}$$

- Observers assign redshift entirely to cosmic expansion and thus infer too low a_{app} at emission
- Cosmological parameters affected by time evolution of

$$\Delta a \equiv a - a_{\text{app}}$$

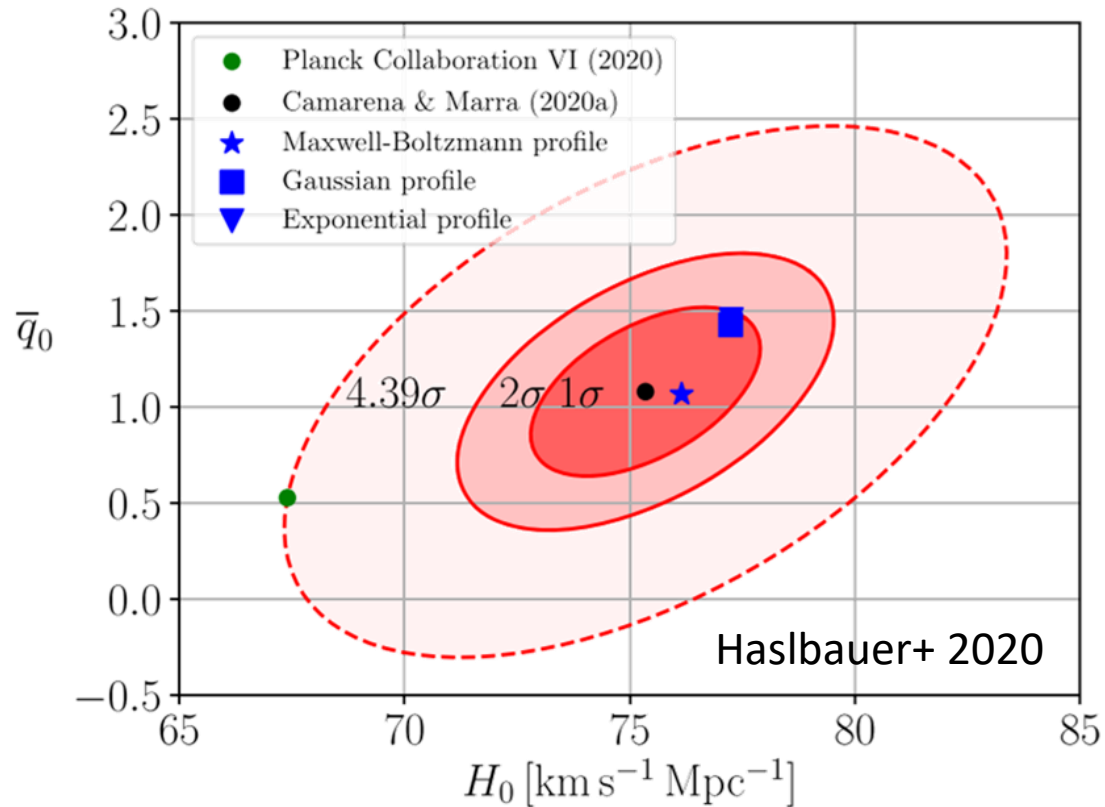
- Biased inference of 1st & 2nd time derivatives of $a(t)$, with obvious cosmological implications.

$$\Delta a(t) = \Delta \dot{a}(t_0) (t - t_0) + \frac{1}{2} \Delta \ddot{a}(t_0) (t - t_0)^2 + \mathcal{O}(t - t_0)^3$$

Local Hubble diagram at second order

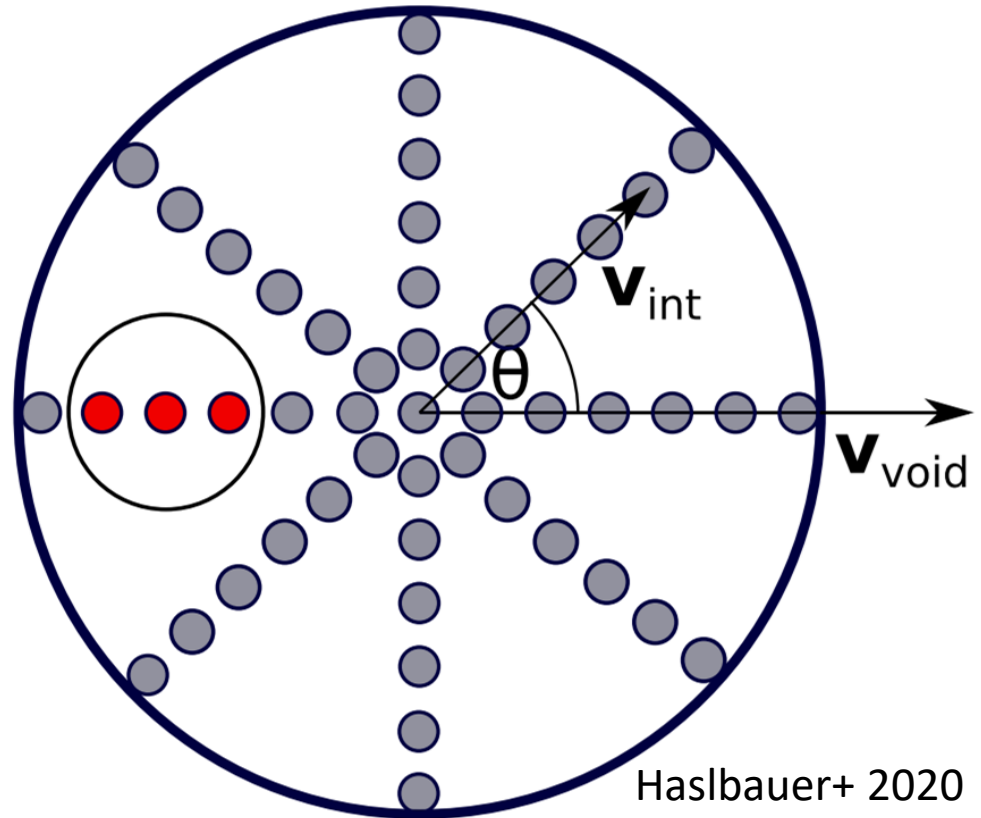
$$H_0 \equiv \frac{\dot{a}}{a}$$
$$\bar{q}_0 \equiv \frac{a\ddot{a}}{\dot{a}^2}$$

- Local observations get these using supernovae with $z = 0.023 - 0.15$ and second order Taylor expansion in $a(t)$
- Model predictions use quadratic regression on time evolution of Δa , the difference at fixed lookback time between **actual** and **apparent** a
- This is added to actual \dot{a} and \ddot{a} in model, which assumes Planck cosmology.



Spherical outflow + systemic velocity

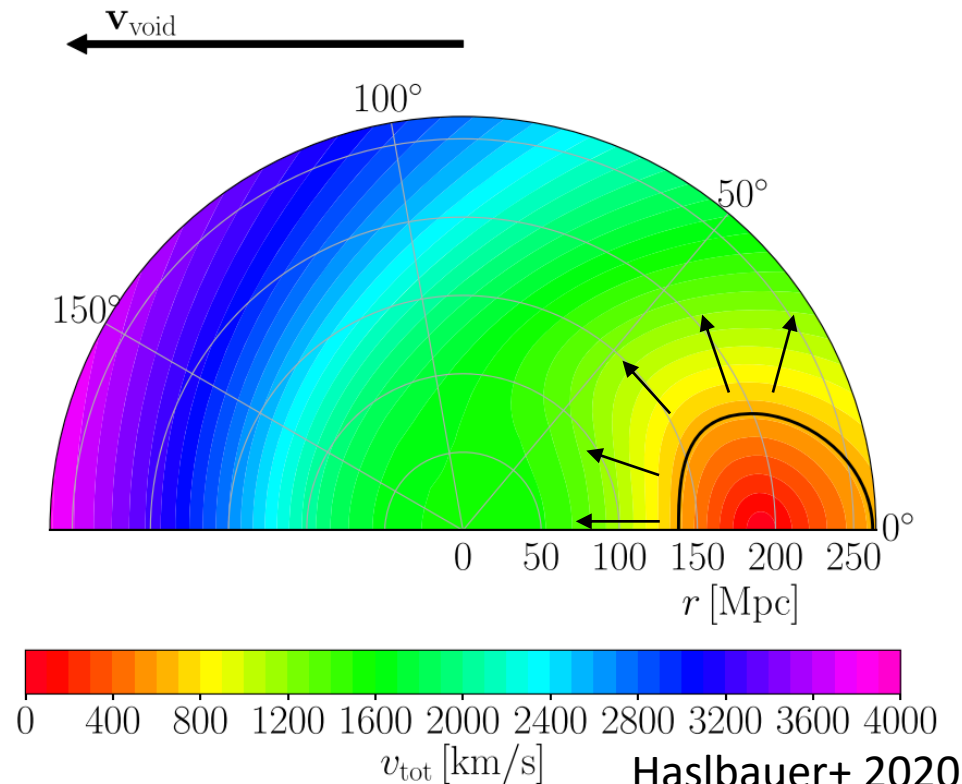
- Model has spherically symmetric outflow + systemic void velocity due to external field
- Void divided into cells
- CMB-frame velocity v_{tot} found in each cell using cosine rule
- Red cells illustrate parts of the void consistent with observed Local Group (LG) peculiar velocity.



Haslbauer+ 2020

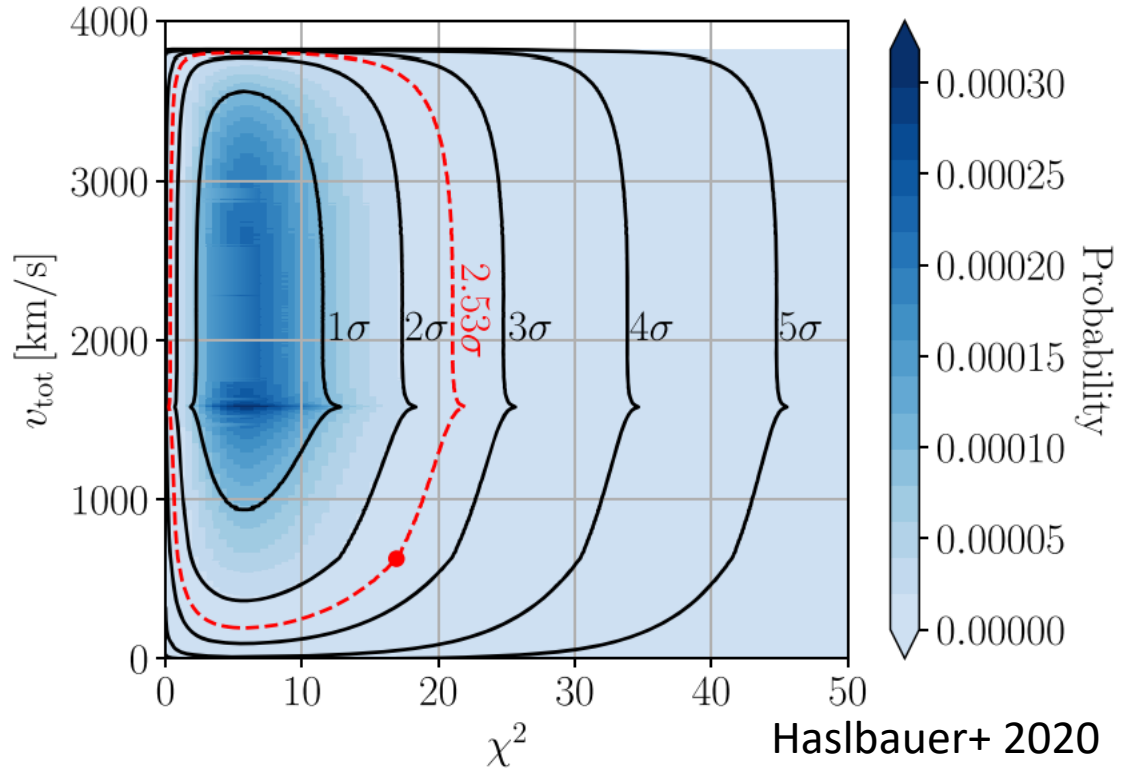
Is our location in the void fine-tuned?

- Colour scale: model velocity in CMB frame (velocities axisymmetric about x-axis).
- Void velocity is time integral of a^*g_{ext} (factor of a accounts for Hubble drag)
- Velocity typically $\gg v_{\text{LG}} = 627 \pm 22$ km/s (Kogut+ 1993), but about 2% of void volume has slower velocity
- Estimated tension is 2.3σ
- We must be close to solid black contour on figure at 627 km/s.



Overall goodness of fit

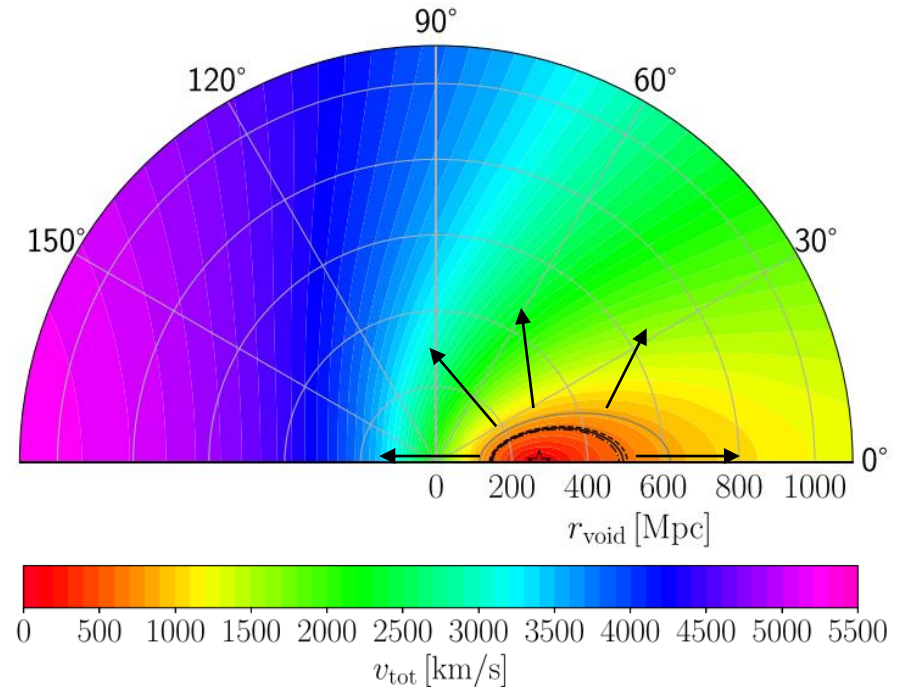
- Model constrained by:
 - local Hubble diagram
 - galaxy number counts
 - observed LG peculiar velocity
- 3 degrees of freedom assumed:
 - initial void size & strength
 - external field
- LG velocity constraint handled separately: CMB-frame velocity on y -axis as its distribution is not Gaussian
- ❖ Overall goodness of fit equivalent to a tension of 2.53σ (pie chart in appendix).



**Bulk
flow**

Void velocity field

- Can test the void model using the velocity field within a few hundred Mpc
- Main problem is that redshifts are due to cosmic expansion and the Doppler effect
 - Hard to disentangle
- Off-centre location in void inevitable
 - Apparent expansion rate would depend on direction.



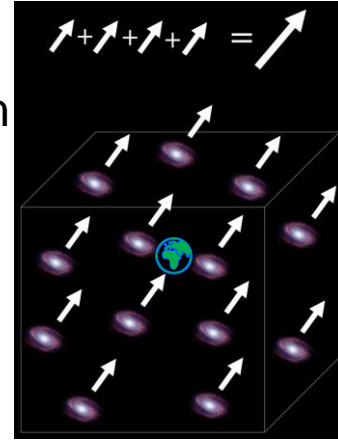
Mazurenko+ (2024)

Bulk flow definition

- Ideally, we want average of 3D peculiar velocity vectors within spherical region
- Need to make do with line of sight peculiar velocity of each galaxy in this region
- Consider this as a vector pointing along the line of sight (LOS)
- Take average of these LOS peculiar velocity vectors, adjust weights so galaxies get higher weights in more sparsely sampled regions (want to sample the velocity field)
- Galaxies weighted by $1/r^2$ (Peery+ 2018).

❖ Consequences:

- If all galaxies have the same peculiar velocity \mathbf{v} , then $v_{\text{bulk}} = v/3$ due to projection effects
- Assumed H_0 has no effect on v_{bulk} as it affects peculiar velocities in spherically symmetric manner, which does not affect the vector average.

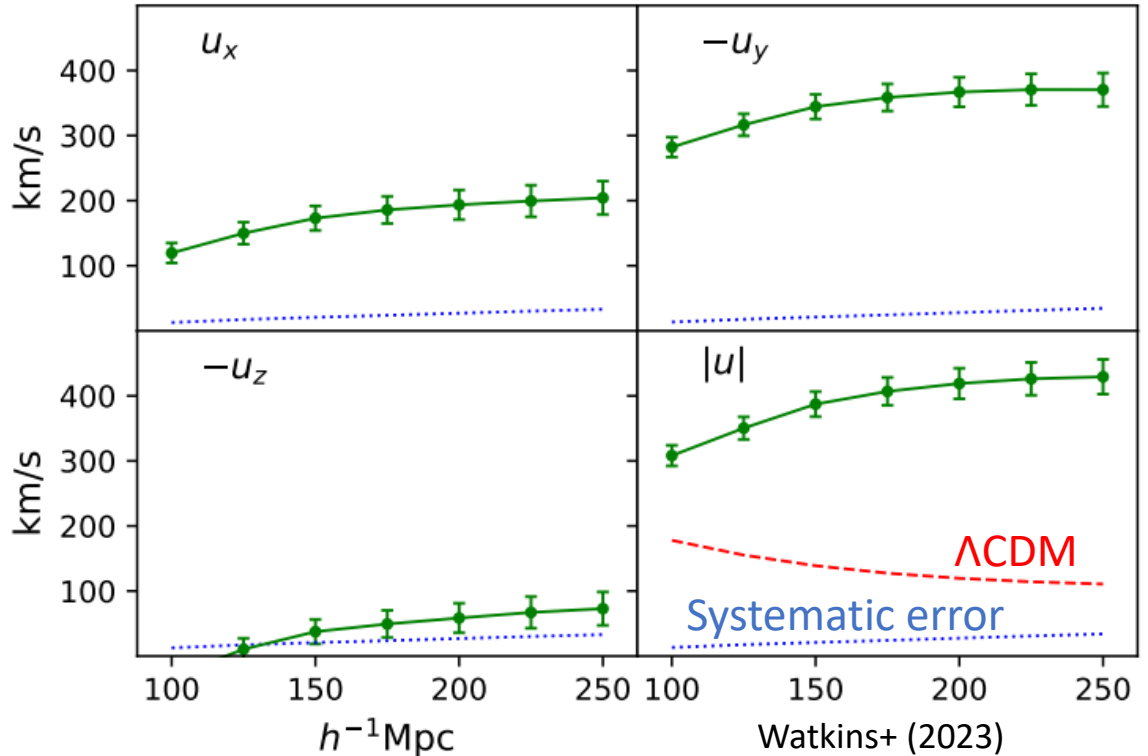


Credit: Abbe Whitford
(from AstroBites)



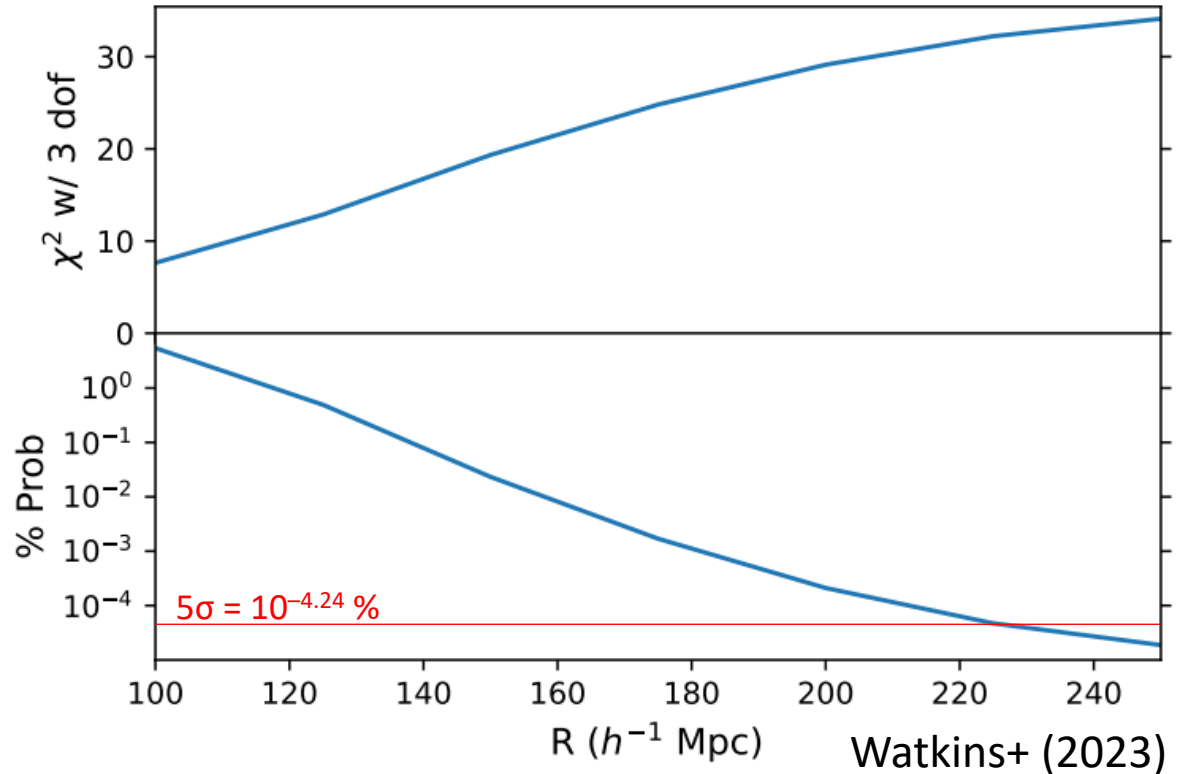
Bulk flow observations

- Need galaxy survey with redshift-independent distances
 - Use CosmicFlows-4 (Tully+ 2023)
- Bulk flow analysis by Watkins+ (2023)
- Subsequent study by independent team (Whitford+ 2023) reports “excellent agreement” out to $173/h$ Mpc, but their method did not extend further out
- The observed bulk flow is independent of the assumed H_0 because changing this affects peculiar velocities in a spherically symmetric manner, not affecting the vector average
- Bulk flow tension \neq Hubble tension.



Bulk flow tension

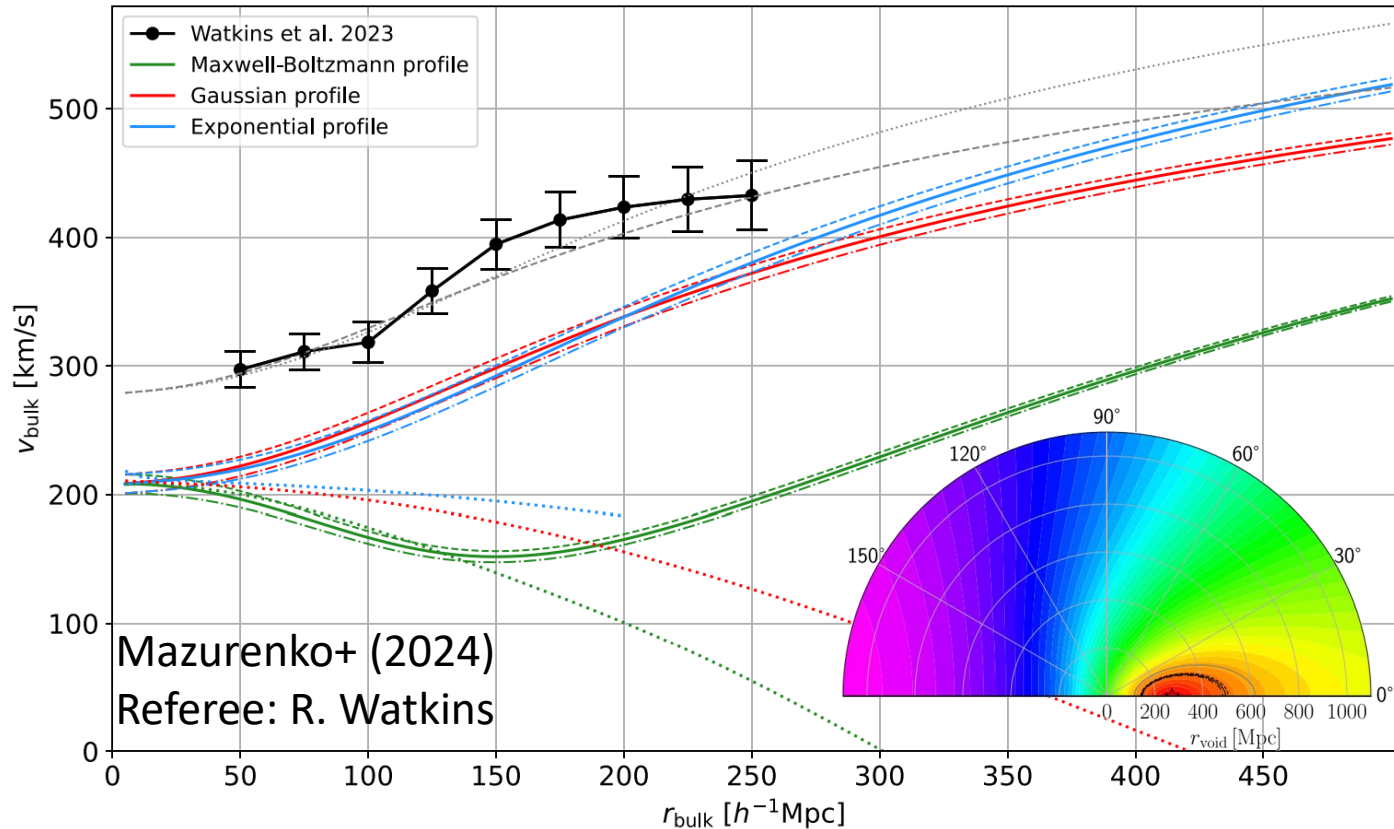
- Expected bulk flow along any direction follows Gaussian distribution
 - Simple χ^2 analysis with 3 d.o.f.
- Beyond $230/h$ Mpc, tension $> 5\sigma$
- Bulk flow tension cannot be solved by adjusting expansion history
 - Problem lies with growth of structure
- No evidence of problems with Λ CDM in CMB on scale of 300 cMpc, which is about twice the first acoustic peak/BAO scale: well measured.



Bulk flow results

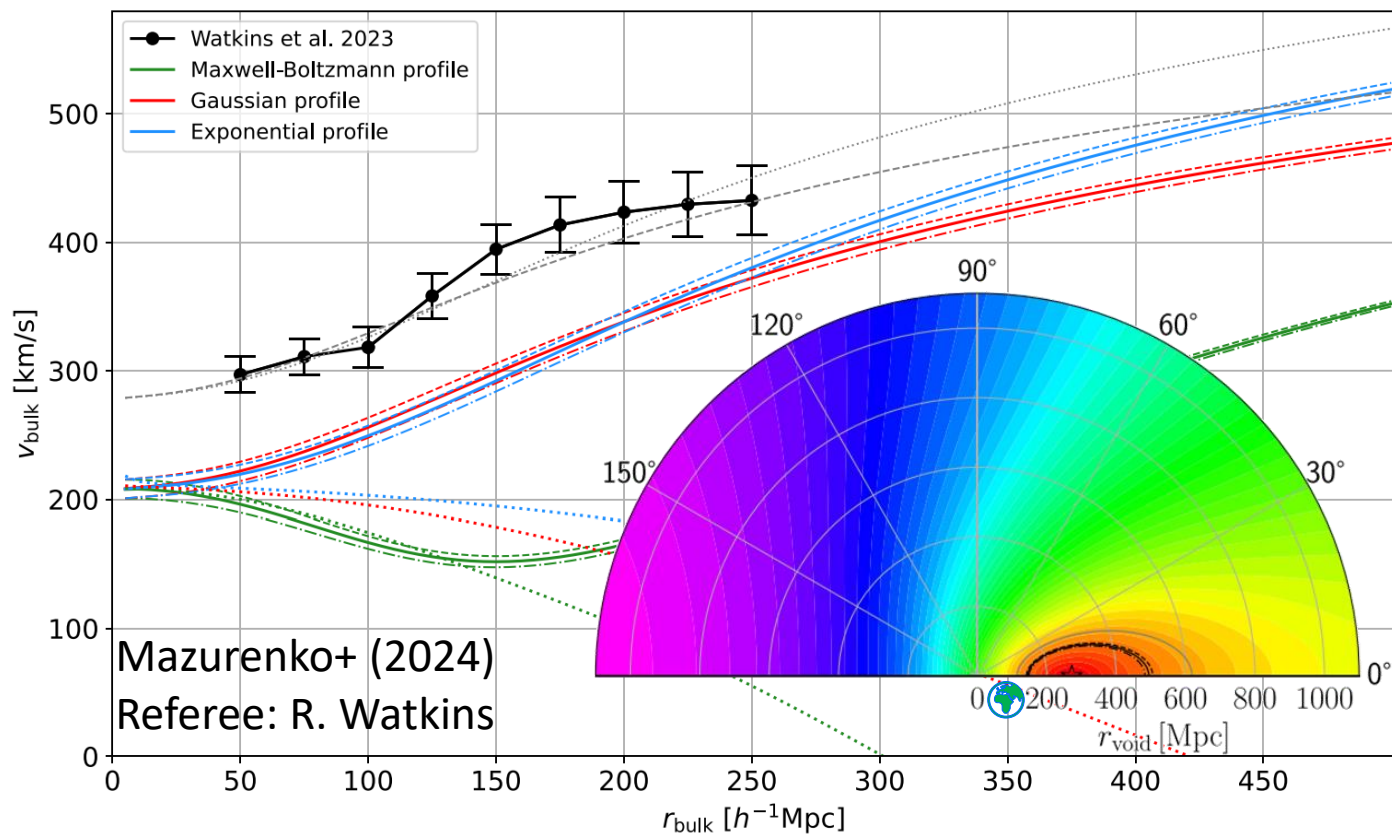
Legend

- Observations (Watkins+ 2023)
- Different colour = different void density profile
- Downward dotted = outer vantage point
- Solid = inner vantage point
- Slightly above/below them = shifts to v_{LG} by ± 22 km/s, which shifts our vantage point within the void
- ❖ Velocities are in CMB frame.



Adjusting our location

- Local Group velocity affected by motions on small scales (nearby galaxies/clusters)
- Ideally, we want the average velocity on a scale larger than galactic separations but smaller than the KBC void
 - Set our vantage point using bulk flow within $50/h$ Mpc
 - Grey curves show results if $v_{LG} = 840$ km/s (dashed = Gaussian, dotted = exp).
- ❖ We are close to the void centre and the void is deepest at its centre.

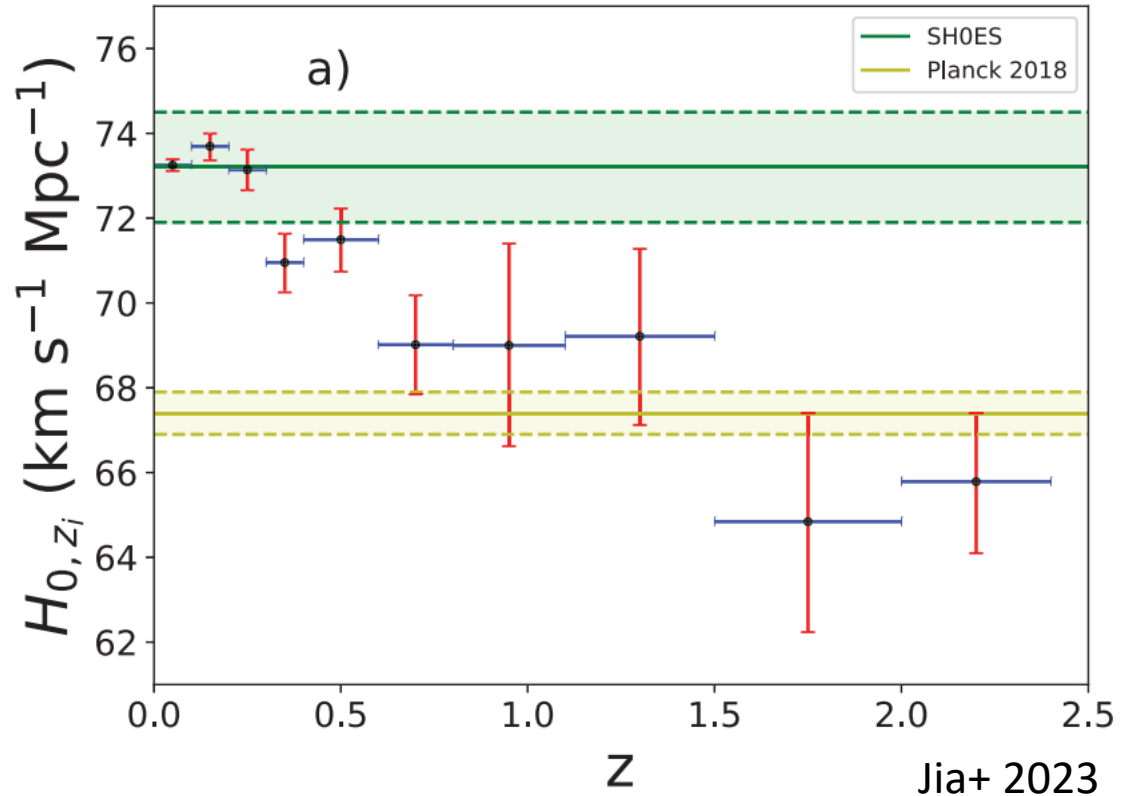


Conclusions

- Galaxy number counts show that we are in an underdensity out to ≈ 300 Mpc or more (Keenan+ 2013, Wong+ 2022) – good evidence across the whole electromagnetic spectrum (see introduction to Haslbauer+ 2020); significant tension with Λ CDM
- Given the nearly uniform initial conditions in the CMB, this requires significant outflow. Generic arguments indicate that this would enhance the apparent local H_0 by $\approx 11\%$
- Locally measured H_0 is larger than predicted in Λ CDM by $\approx 10\%$ (di Valentino+ 2021)
 - Hubble & void tensions were linked in semi-analytic void model of Haslbauer+ 2020
- Predicted bulk flow curve was later observed (Mazurenko+ 2024; Watkins+ 2023)
- Measurements at $z > 0.5$ or so are consistent with background Planck cosmology
- Early time solutions to the Hubble tension face at least seven difficulties (Vagnozzi+ 2023)
- Very few proposed solutions to the Hubble tension had a different motivation and made a successful *a priori* prediction of a different phenomenon that is unlikely in Λ CDM.

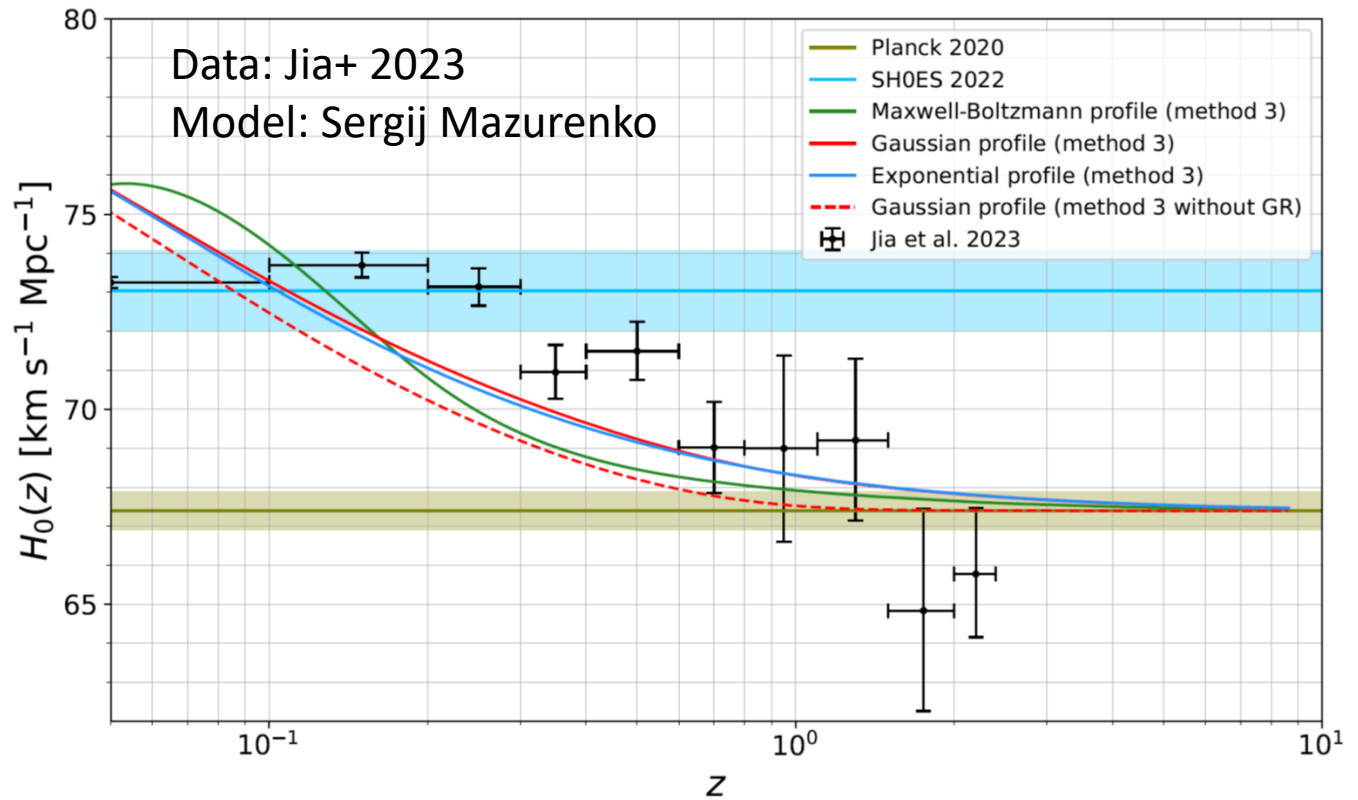
Hubble tension at high z ?

- Model has large local void with background Planck cosmology
- Observations at high z must recover Planck cosmology as redshifts from peculiar velocities would be small compared to cosmic redshift
- Need to carefully infer H_0 from only the data within a narrow redshift range
- Such analyses do show the expected return to Planck cosmology
- No Hubble tension at high z
- Systematic calibration error would inflate all H_0 values by common factor.



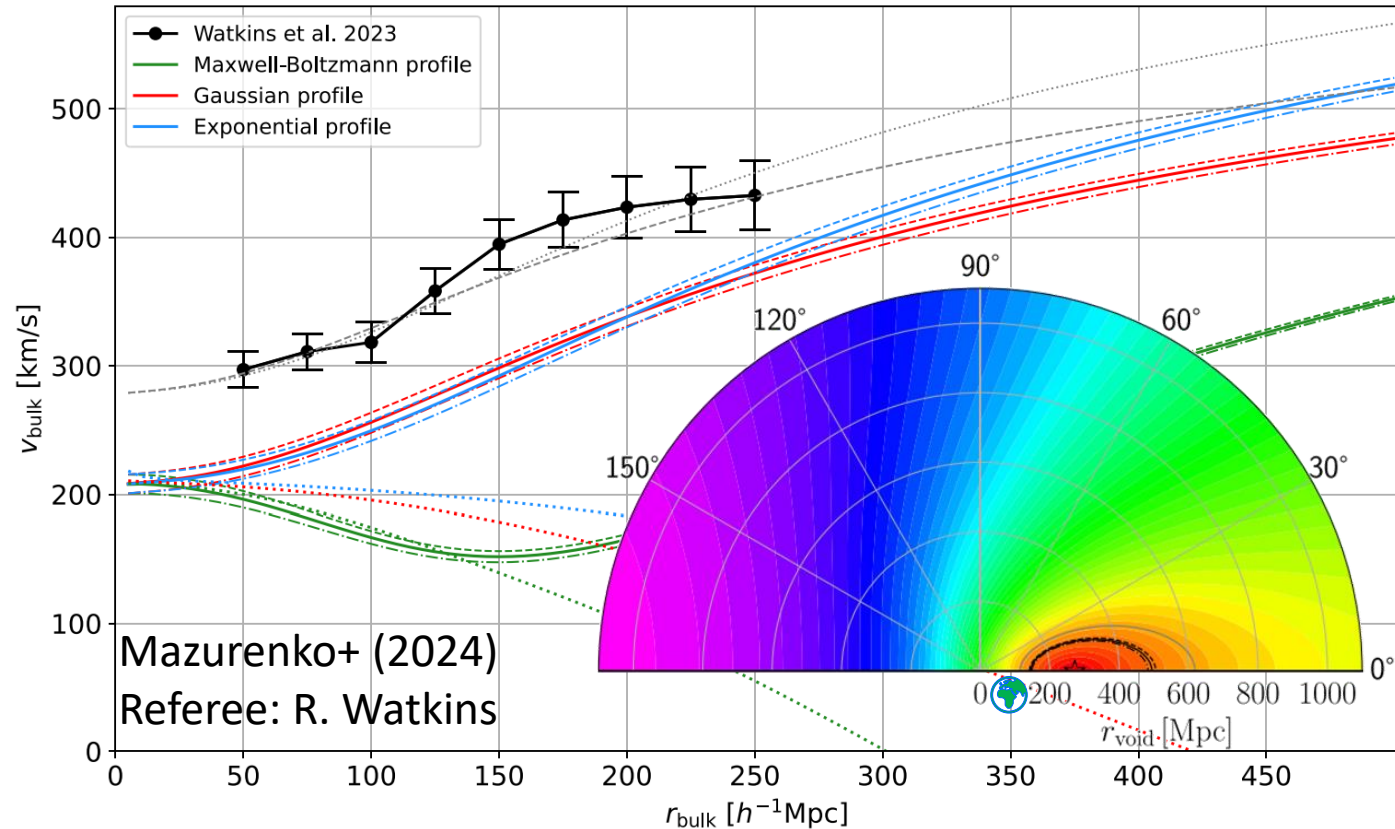
$H_0(z)$ in local void scenario

- Estimated H_0 that would be inferred by an observer from data in narrow redshift range if the local void model is correct
- Idea is to suppose observations constrain lookback time independently of z , e.g. from luminosity distance
- z then used to infer a at that time assuming only cosmic expansion causes redshift (wrong in this scenario)
- Data then fit with FRW model.



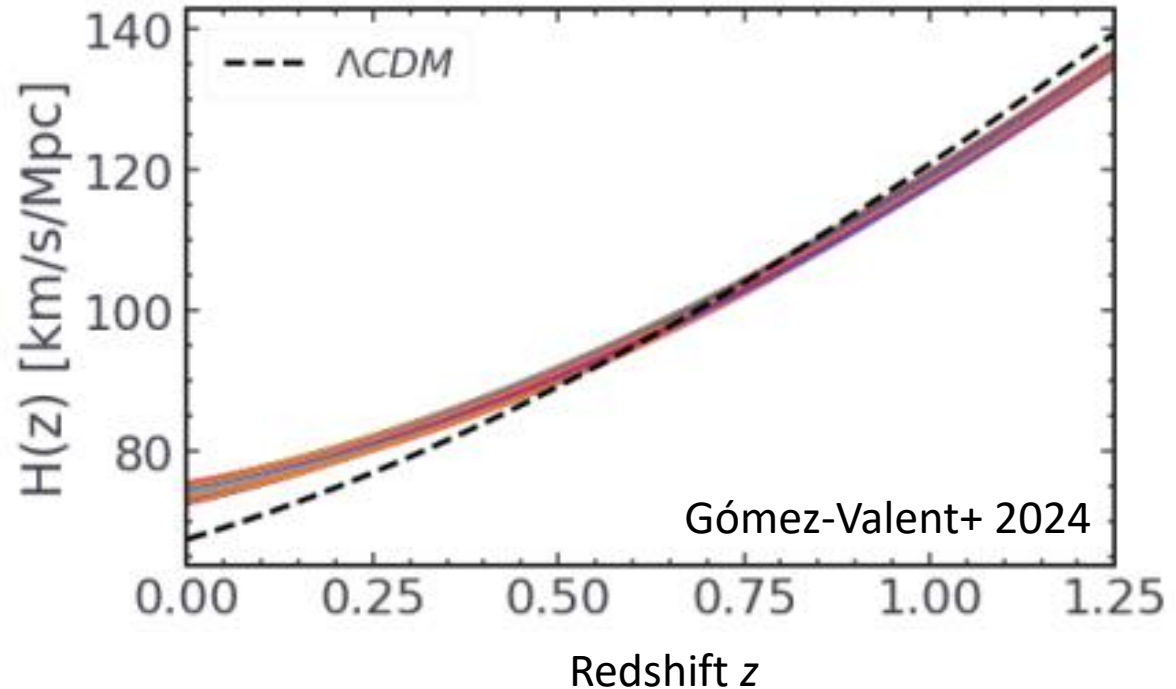
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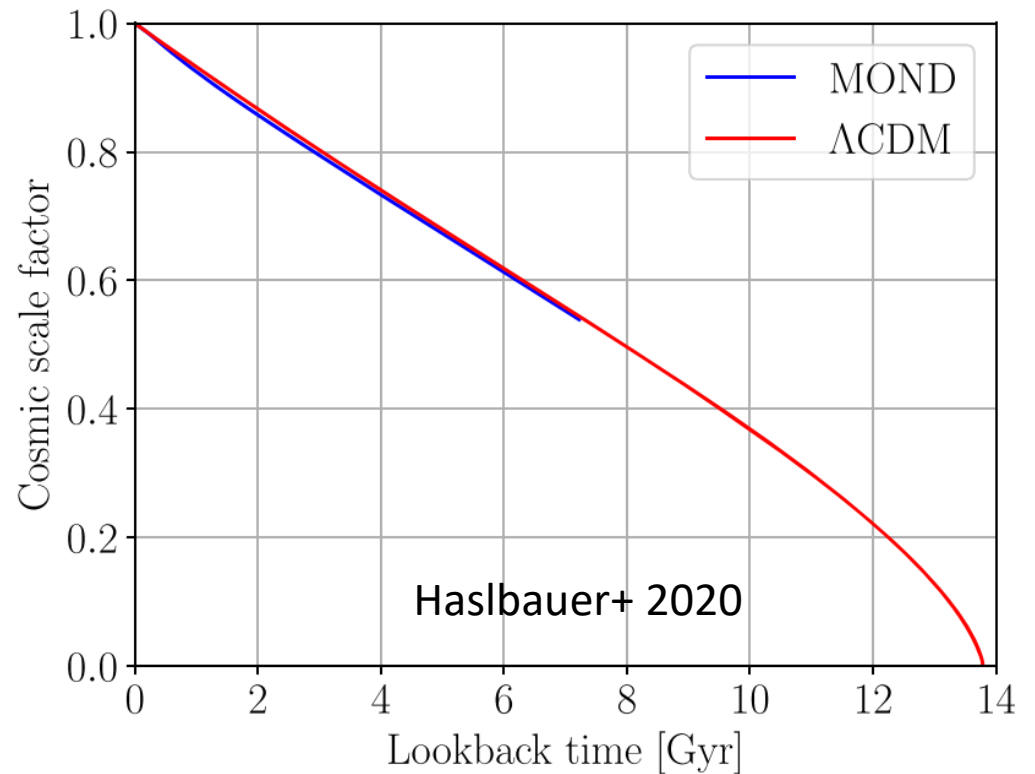
Reconstructed Hubble diagram

- Can infer expansion history $H(z)$ without assuming any model
- This seems to follow the Λ CDM expectation until recently
- Small compensatory adjustment at high redshift to preserve angular diameter distance to CMB, but there is anyway some uncertainty in the official Planck parameters
- Combined with other difficulties faced by early time solutions to the Hubble tension (Vagnozzi+ 2023; Cimatti & Moresco 2023), recent results favour a more recent or local solution.



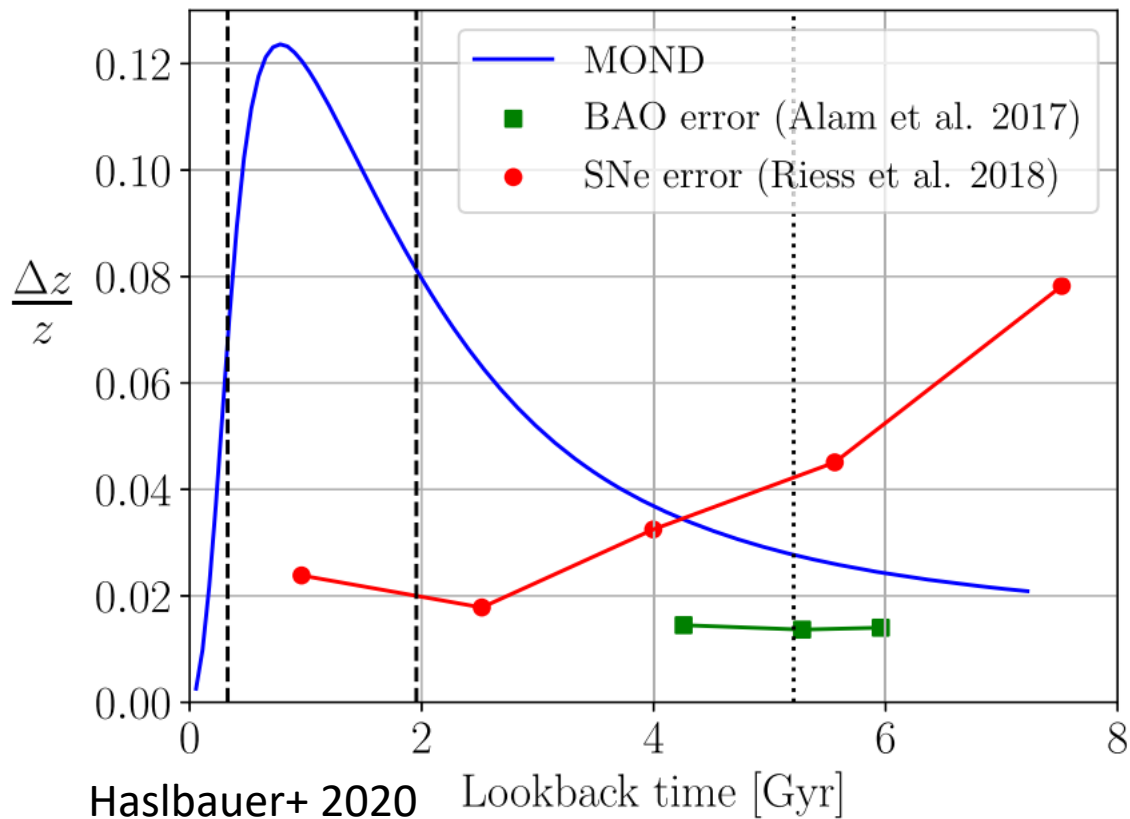
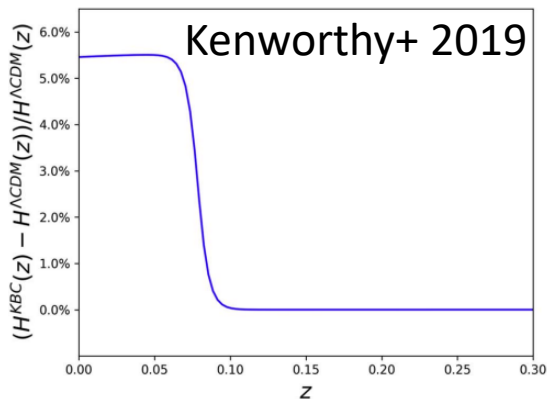
A subtle change can solve the H_0 tension

- A subtle difference between the actual and apparent expansion histories can solve the Hubble tension.
- Local void solutions create peculiar velocities out to quite high redshift in MOND due to slower decay of the gravity law
- This means the return to a Planck cosmology could be fairly slow, contradicting the assumptions in the famous Kenworthy+ 2019 paper which argued against the directly observed local void (Planck cosmology assumed to be fully recovered when $z > 0.1$).



The void distorts redshifts quite far out

- Blue curve in figure on right shows that void enhances redshifts about 10% in the range typically used to obtain the local H_0 (between the dashed vertical lines showing $z = 0.023 - 0.15$)
- Notice that peculiar velocity corrections are important quite far out.



Overall goodness of fit of local void model

- Overall tension is 2.53σ
- Pie chart: summary of individual contributions
- v_{LG} tension is based on fraction of void volume with a slower velocity in CMB frame
- Bulk flows not considered
- ❖ No substantial tensions.

(Degrees of freedom, equivalent tension for 1D Gaussian)

